## Gamma Ray Observatory to April 5, 1991



## Early Capabilities Limited

- Limited or no telemetry
- Tape recorders (no solid state memory)
- **Discrete electronics**
- Limited processing power
  - In space
  - on ground
- Limited position and timing capability needed for pulsar studies
- Limited payload capacity for space launches

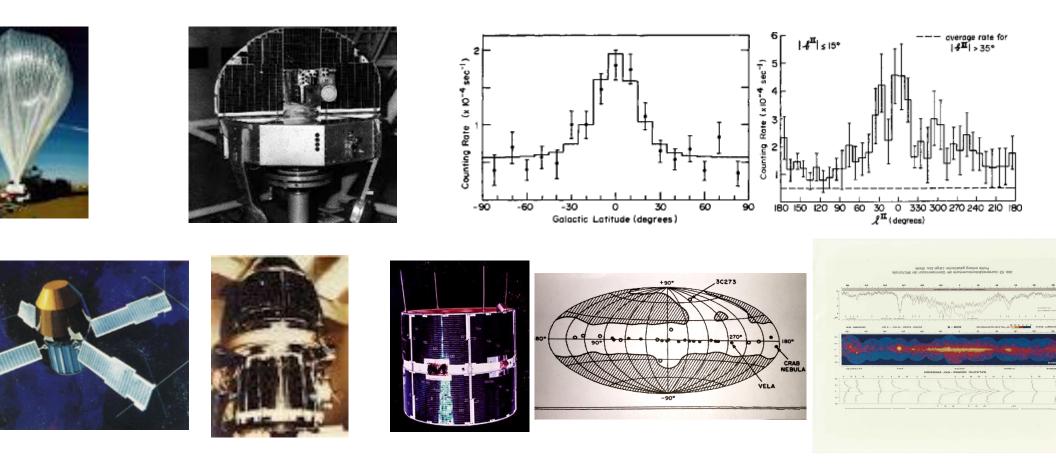
## Motivation for Gamma-Ray Astronomy

- Are cosmic rays galactic or extra-galactic?
- If cosmic rays are galactic how can you probe the galactic distribution?
- If galactic what is the distribution of protons versus electrons?
- Are the newly discovered gamma-ray bursts galactic or extragalactic in origin?
- Detectors not traditional telescopes as in other portions of the electromagnetic spectrum by detectors borrowed from the physics community.

## Theoretical Motivation

- Feenberg and Primakoff, 1946, "Interaction of Cosmic-Ray Primaries with Sunlight and Starlight" Phys. Rev. Letters.
- Fermi, 1955, "Nuclear Physics", Notes Compiled by J. Orear.
- Morrison, 1958, "On Gamma-Ray Astronomy", Il Nuovo Cimento.
- Ginzberg and Syrovatskii, 1965, "Some Problems of Gamma- and X-Ray Astronomy, Soviet Phys.-Uspekhi
- Fazio, 1967, Gamma Radiation from Celestial Objects, Annual Rev. of Astronomy and Astrophysics.
- Discoveries not based on theory: Gamma-ray Bursts (Vela), Gamma-rays from AGN (EGRET), Gamma-Ray Pulsars

## Before GRO



# Pre-GRO Launch Operations



Launch Operations of NASA's second Small Astronomy Satellite from the Italian San Marco launch platform in the Indian Ocean off the coast of Kenya in 1971.

## Gamma Ray Observatory



November 15, 1977

llowing the premature ending to SAS-2, Fichtel and Macdonald worked tirelessly to convince NASA to include servatory to explore the full gamma-ray portion of the electromagnetic spectrum. Originally announced as a ar mission when launched it had become a three year mission that was deorbited after nine years on orbit.

#### WHO IS INVOLVED IN THE GAMMA RAY OBSERVATORY?

**GODDARD SPACE FLIGHT CENTER GRUMMAN AEROSPACE CORPORATION** MARSHALL SPACE FLIGHT CENTER NAVAL RESEARCH LABORATORY UNIVERSITY OF NEW HAMPSHIRE NORTHWESTERN UNIVERSITY **RICE UNIVERSITY** STANFORD UNIVERSITY ESTEC DEPARTMENT OF SPACE SCIENCE, NETHERLANDS UNIVERSITY OF LEIDEN, NETHERLANDS MAX-PLANCK INSTITUT, GARCHING, WEST GERMANY **ROYAL AIRCRAFT ESTABLISHMENT** 

### WHY THE GAMMA RAY OBSERVATORY?

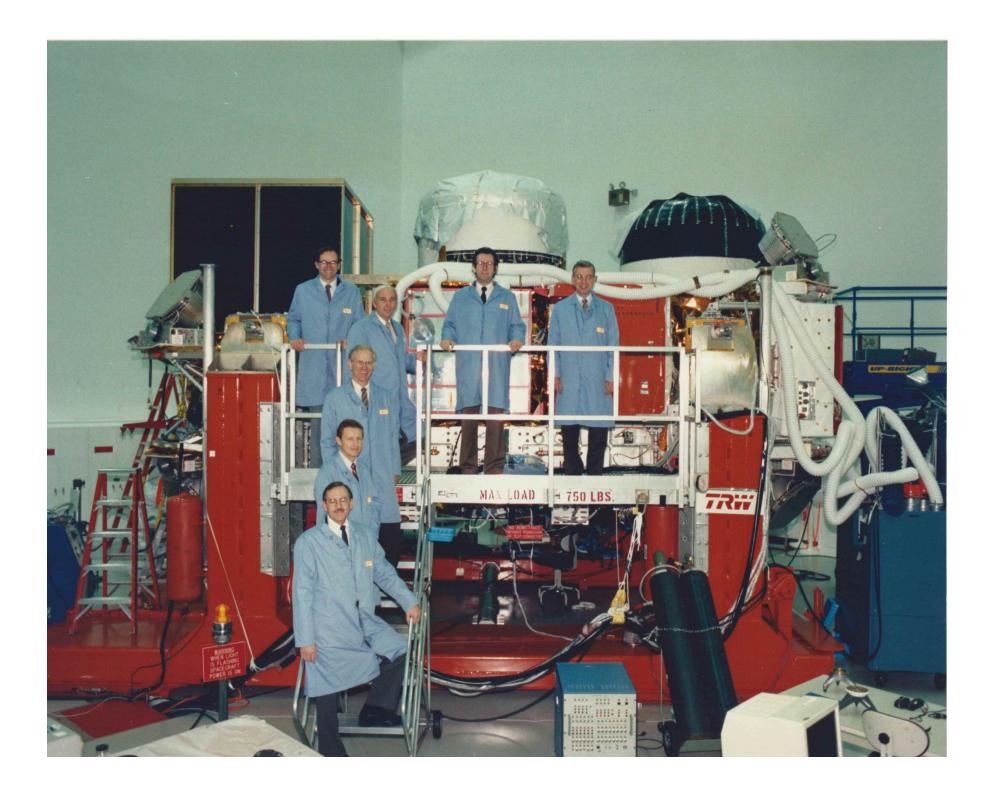
 FIRST MISSION TO STUDY THE FULL RANGE (OVER FIVE DECADES) OF GAMMA RAY PHENOMENA

 WILL SEE SOURCES 10 TIMES FAINTER THAN PREVIOUS INSTRUMENTS

• WILL IMPROVE KNOWLEDGE OF SOURCE POSITIONS BY FACTOR OF 50-100

 WILL EXPLORE PREVIOUSLY UNOBSERVED PORTIONS OF GAMMA RAY SPECTRUM

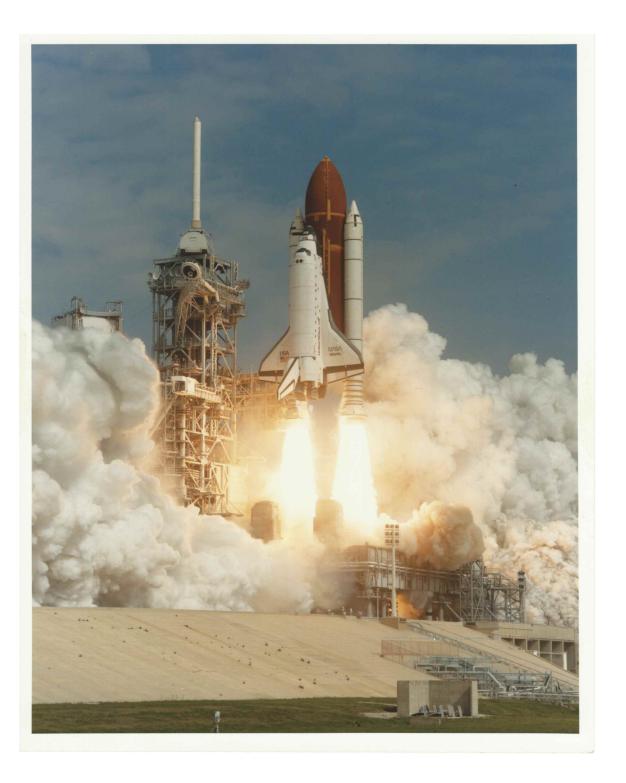












April 5, 1991 09:22:44 EST after 4:45 minute weather hold

Heaviest scientific p (35,000 pounds), ev at the time.

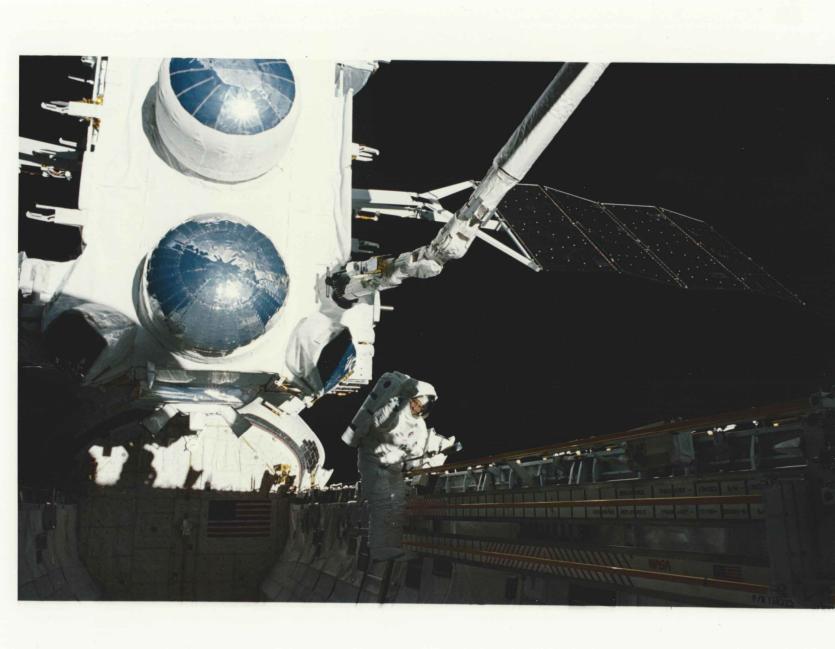


Launched on a rainy morning on April 5, 1 Caught the User's Cor by surprise while they looking the other way



\$37-51-021

Lyndon B. Johnson Space Center Houston, Texas 77058



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## What did we learn?

- Extended mission duration led to a robust Guest Observer program
  - Run by the CGRO Science Support Center the CGRO GO was well managed and was successful in generating many important user investigations.
  - GO Participation was difficult to support because the GO program was an after thought when the program was extended to 3 years and beyond. Each instrument had its own data processing and archiving system.
  - With a small gamma-ray astronomy community, the CGRO GOs became a large and vocal support group not just for CGRO, but built community support for a follow-on mission (Fermi Gamma-Ray Observatory).
- The Compton Fellowship Program was very successful in attracting bright productive young scientists into the field.
- The close working relationship between the GRO Project Staff and the PI Teams led to many enhancements in the scientific return. The mission might have failed at the start had not the GRO Project Staff planned for the possibility of problems in the deployment of the appendages (solar panels and high-gain antenna). The CGRO scientific community will forever be indebted to the heroic efforts of Jerry Ross and Jay Apt in rescuing what could have been a very crippled Observatory. This was the first extravehicular activity in 6 years.
- The most significant early failure of both tape recorders actually was a fortuitous benefit to the study
  of gamma-ray bursts allowing immediate notification of bursts to other observers for followup
  observations.