

lcplot

June 2, 2019

Abstract

lcplot produces plots of the net source and background time series for a single time series FITS file or for several time series FITS files produced by elcbuild for EPIC and omlcbuild for the OM. Furthermore it performs either the Chi Squared probability of constancy test and/or the Kolmogorov-Smirnov probability of constancy test on the light curve to determine whether the source is variable.

1 Instruments/Modes

Instrument	Mode
EPIC MOS	IMAGING, TIMING
EPIC PN	IMAGING, TIMING
OM	FAST

2 Use

pipeline processing	yes	
interactive analysis	yes	

3 Description

lcplot produces plots of time series measured either by EPIC or by the OM. These plots show both net and background count rates. **lcplot** can be used to display a single time series directly or the multiple time series, produced by **elcbuild** for EPIC and by **omlcbuild** for the OM, and supplied as FITS files. Alternatively the task **omthlcplot** can be used to produce rate plots of the OM tracking star's intensities.

The input FITS files must contain net count rates, background count rates and the associated errors. The time series must follow a regular binning scheme (i.e. equispaced time bins) [1]. The bin width is given by the keyword TIMEDEL.

Each light curve (net and background) of the input file is plotted on a separate page. For instance, in a PPS use, EPIC time series are given in three energy bands, so the plots will be composed of three light curve pages. On the other hand, an OM FITS file which contains only one light curve, will give a single page plot.

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The task performs 1) the Chi Squared probability of constancy and/or 2) the Kolmogorov-Smirnov probability of constancy tests on the total counts. **Icplot** tests the null hypothesis (the source is not variable) using 1). A Chi Square fit is carried out on the count distribution and the constant distribution whose value is equal to the observed mean (both statistics and the probability are computed) using 2). A Kolmogorov-Smirnov test is carried out on the cumulative probability functions of the observed count distribution and the theoretical Gaussian distribution whose mean and variance are both equal to the observed mean count.

Results are written on a single page after the light curve(s).

Plots give time after TSTART (keyword in the FITS file) on the X-axis against count rates on the Y-axis. Please note, bins with a low exposure and therefore large errors are not plotted.

lcplot allows the user to input the device type (e.g. Xwindow, Postscript, ...) for output. The y-scale of the background counts can be forced to be the same as the source counts by setting the parameter 'bkgdyscale' to 'yes'. Otherwise the y-axis is scaled to the most appropriate value for the background counts.

As described in [1], light curves should be rebinned if necessary in order to limit the number of plotted data points to 500.

3.1 Examples

To create a postscript plot, called plot.ps, with both net and background count rates and the results of the Chi Square and Kolmogorov-Smirnov tests, from the FITS file 'lightcurve.ds', where the counts are binned up by a factor 10, a command such as,

lcplot set="lightcurve.ds" plotdevice="/VPS" plotfile="plot.ps"
binsize=10 tests=both

can be used.

4 Parameters

This section documents the parameters recognized by this task (if any).

Parameter	Mand	Type	Default	Constraints

set	yes	dataset	No default			
Name of the FITS file containing the time series. Case sensitive						

Name of the FTTS file containing the time series. Case sensitive.

plotdevice	no	String	"/PS"	Constraint : valid de-
				vice type

Not case sensitive.

plotfile	no	String	"plot.ps"	less than 1024 char.
Name of the output file conta	ining the sl	ot Conson	aitima	

Name of the output file containing the plot. Case sensitive.

$\mathbf{binsize}$	no	Integer	5	Constraint : positive
				value

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VOC	/no

bkgayscaie	no	boolean	no		yes/no	
To scale the y-axis of the plot	of the back	ground cou	nts to be the s	same as the	y-axis of the	plot of the
source counts						

tests	no	string	both	both/chi/ks

Should both (both) tests be carried out or simply the chi-squared (chi) or the Kolmogorov-Smirnov (KS) test

5 Errors

This section documents warnings and errors generated by this task (if any). Note that warnings and errors can also be generated in the SAS infrastructure libraries, in which case they would not be documented here. Refer to the index of all errors and warnings available in the HTML version of the SAS documentation.

Incorrect RATE extension (error)

Incorrect RATE extensions (see Developer's notes for more details)

Keyword TIMEDEL missing (error)

The keyword TIMEDEL is missing

Existing output file cannot be overwritten (error)

The output file exists, and cannot be overwitten because the SAS_CLOBBER variable is set to 0

Unavailable default device type and/or output file (error)

Unavailable default device type and/or output file

Invalid input file (error)

Invalid input column dimensions (the four columns RATE, ERROR, BACKV, BACKE contained in the FITS file must be arrays of dimension Nlightcurve \times Nbins)

TestTypeUnknown (error)

The tests available are 'both', 'chi' or 'ks'

Unavailable device type and/or output file (warning)

corrective action: The device type and output file are set to the defaults.

Number of data points to plot is higher than 500 (warning)

 $corrective\ action:$ Increase binsize value until the number of points to plot is lower than 500 [1]

Too small number of bins (warning)

Number of valid bins is lower than 2. Cannot perform tests corrective action: No variability tests are delivered



Mandatory keywords are missing from the input file (warning)

Some mandatory header keywords are missing or inconsistent corrective action: A warning is raised and the task tries to proceed

No valid data points (warning)

All the input values are NaN corrective action: A warning is raised, the NaN values are set to 0.0 and the task continues

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6 Input Files

1. EPIC or OM FITS Time Series [1] (generated by **omlcbuild** or **elcbuild**).

7 Output Files

1. EPIC or OM Postscript Source Time Series [1] (Deliverable PPS product)

8 Algorithm

```
Recover parameters (File names, Device type, Binning factor, Ouput file)

Read source time series FITS file:
    Get dataset and table.
    Check important keyword consistency (notably TSTART and TIMEDEL).
    Call a warning or error if necessary.

Recover all light curves included in table (in PPS use, 3 for EPIC, 1 for OM):
    Net source rates and errors & background rates and errors are recorded in arrays of dimension Nlightcurve * Nbins.

Open plot: create view surface in specified device format (check its validity).

For each light curve:

Delete gaps in data (when the IEEE NaN constant is found).

Increase binsize if necessary to limit the number of points to 500.

Rebin rate and error columns by binsize

Create a labelled plot page separated into two panels in which time is graduated
```

on X-axis (see Comments), and count rates on Y-axis (in counts/sec).

Draw net source and background time series on different panels.

Reconstruct net source counts and total counts rebinned :

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Perform variability tests on rebinned net source counts : Calculate mean count rate and variance.

Test the null hypothesis (the source is not variable) by :

- Processing a Chi Square fit between the count distribution and a constant distribution whose value is equal to the observed mean (statistic and probability are computed).
- Processing a Kolmogorov-Smirnov test between the cumulative probability functions of the observed count distribution and a theorical Gaussian distribution whose mean and variance are both equal to the observed mean count.

End for

Write variability test results on a single plot page.

Release memory.

Close plot.

End subroutine lcplot

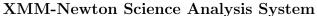
9 Comments

- **lcplot** only produces plots of light curves measured by EPIC or OM. We note that no RGS Source Time Series FITS file is specified in [1].
- Error rates are rebinned following a quadratic scheme (Central limit theorem).
- Time intervals all have the same length, the first one is [TSTART, TSTART + Binsize*TIMEDEL] where TIMEDEL is the bin width read from the file [1].
- We note in [1] that the keyword TIMEDEL is specified in MJD. We have assumed it to be given in seconds.
- Time on the X-axis is given as T-TSTART in seconds.
- The mean, variance, Chi Square, etc are provided with six significant digits.
- A Chi Square statistic close to the bin number and a small Kolmogorov-Smirnov statistic with probabilities higher than 0.01 indicate that the source is not variable.
- The output devices are those of the PGPLOT library. The device list is governed by the drivers available to PGPLOT.

10 Future developments

- A parameter Time_Unit could be added to the parameter file. It would allow a choice of different time axis units (e.g. Hours after TSTART).
- The results of the statistical analysis could be saved as keywords in the header of the light curves source FITS file. Example, for the first energy band, one could add:

AVRGE_1 = / Avg, count/s in light-curve



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 $FREXP_1 =$ / Avg. Frac exposure in light-curve VAROB_1 = / observed variance VAREX_1 = / expected variance THRDM_1 = / third moment $MININ_1 =$ / Minimum Intensity / Maximum Intensity $MAXIN_1 =$ EXVAR_1 = / excess of variance $CHI2_1 =$ / Chi squared / Chi squared probability CHI2PR_1= / K-S test value KS_1 / K-S probability $KSPR_1 =$ $RMS_1 =$ / RMS variability AVRGE_E1= / Error Avg count/s VAROB_E1= / Error observed variance / Error expected variance VAREX_E1= $CHI2_E1 =$ / Error Chi2 / Error RMS variability $RMS_E1 =$

10.1 CAL usage

• None

References

[1] SSC. XMM Survey Science Centre to Science Operations ICD for SSC Products. Technical Report XMM-SOC-ICD-0006-SSC Issue 2.1, SSC, Mar 2000.