January 19th, 2024 The 2nd XRISM Community Workshop @ UMD

XRISM Study on Ionized Absorbers of MCG-06-30-15

<u>Shoji Ogawa (ISAS/JAXA)</u>

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AGN Structure



AGN Structure

- Supermassive blackhole, Accretion Disk, Torus, Outflow



Revealing structure of AGN is important for understanding AGN feeding/feedback mechanisms



X-ray Spectrum of AGN



X-ray spectrum of an AGN carries information on:

- All material with various temperatures and ionization states
- Line-of-sight materials (absorption feature)
- Torus and/or accretion disk (reflection component)



Aim of This Research



Spectral Complex in Fe-K band:







XRISM/Resolve enables us to separate these component



Torus (disk) reflection accompanied by narrow fluorescence emission lines (Fe Ka@6.4 keV)



Broad Fe Ka



Relativistic reflection from innermost region of the accretion disk













The line profile of Fe Kα observed by the ASCA (Tanaka+95)

- Absorbed Continuum shape can mimic Broad Fe Kα with current CCD energy resolution
- Narrow features associated with ionized absorbers can be easily detected with XRISM/Resolve





Target: MCG-6-30-15



MCG-6-30-15 is an archetypical Seyfert 1 galaxy showing:

- Complex absorption
- Proad Fe Kα emission
- Variability in flux and spectral shape
- Two spectral model (relativistic reflection, Tanaka+95; partial covering, Miyakawa+12)







Joint Observation



- Midooka+22
- Velocity
- \rightarrow Revealing the structure of AGN

- NuSTAR: continuum shape
- Chandra or XMM-Newton
 - Soft X-ray band (LETG, RGS): lower ionized gas
 - UV (XMM-Newton/OM): photoionization continuum

- Complete picture of ionized outflows
 - **Ionization Parameter**



Summary

Aim of this research:

- To detect narrow features associated with ionized outflows and reveal their physical properties
- To separate between these narrow spectral features and the relativistic reflection signals from the accretion disk (if any)
- To correctly determine the innermost disk radius and black-hole spin from the "disk-line" profile

Long-term Observation of MCG-6-30-15

- To perform spectral fit to Intensity-sliced spectral
- To detect absorption features
- To test if partial absorber model is suitable

